



# Math and Global Temperature

Mary Lou Zeeman  
Bowdoin College  
&  
Math and Climate  
Research Network



Thanks to:  
**many friends  
& colleagues**



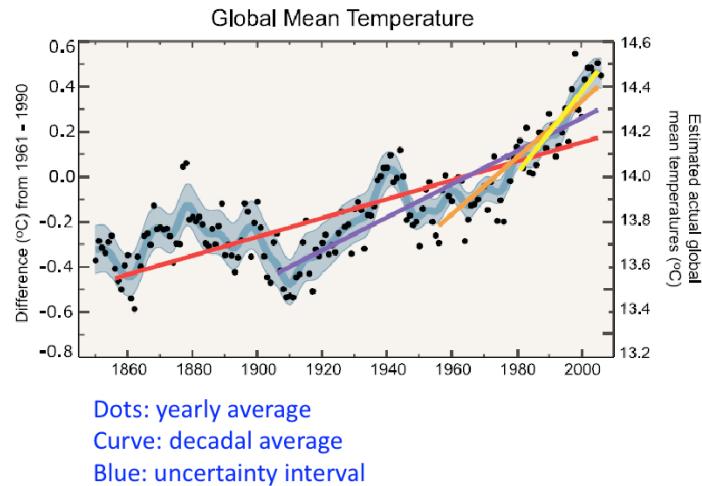
**Bowdoin**

This talk was originally prepared by Mary Lou Zeeman for a lecture to the Coastal Studies for Girls residential high school in Freeport, Maine, in 2010. It has been modified slightly since then but remains the same in spirit. Feel free to use it. Please make sure to credit sources of information and images.



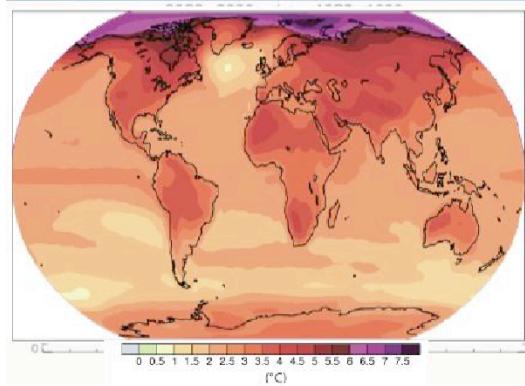
## Evidence of a Changing Climate

Intergovernmental Panel on Climate Change (IPCC):  
Warming is “unequivocal”



We're going to focus on global mean temperature  
(mean over planet and over a year)

## Why do we try to predict?



Warming is NOT uniform

IPCC

Need to make adaptation policy decisions to protect areas of vulnerability

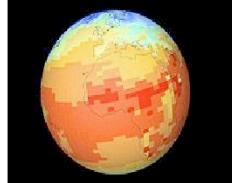


## Why use Mathematical Models?



1

many

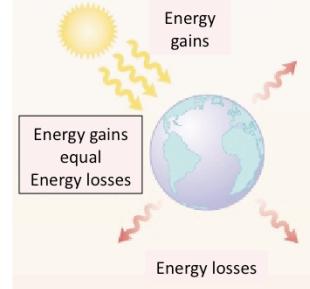


C. Jones

We've only got one real experiment...



## Simplest Energy Balance Model



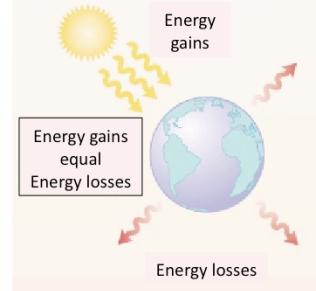
CIMSS/ U Wisc

Simple concept that underlies a family of models



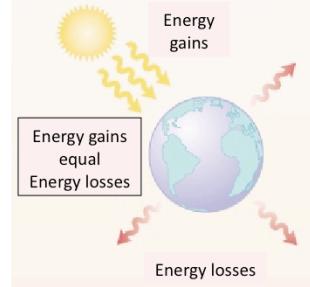
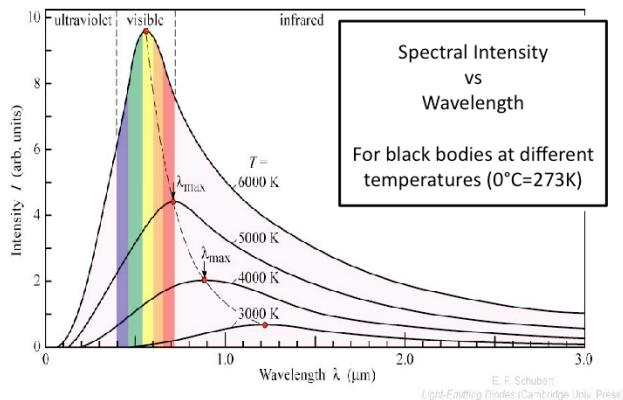
## Simplest Energy Balance Model

**PLAN:** Use this model to understand the role of greenhouse gases in regulating our temperature



# Understanding Radiated Energy

## Black Body Radiation - Planck



Sun: ~6000K, emits mainly in ultraviolet (UV) and visible spectrum (shortwave)

Earth: ~300K, emits mainly in infrared (IR) spectrum (longwave)

Total energy flux over all wavelengths for temp T is  $\sigma T^4 = (5.67 \times 10^{-8}) T^4$  -- Stefan-Boltzmann

A “Black Body” absorbs everything at all wavelengths

It radiates energy according to it’s temperature (discovered by Planck)

Shorter wavelength has higher energy

Earth is an order of magnitude cooler than any of the graphs shown

So Earth radiates energy at much longer wavelengths than sun

Degrees K and C are the same size

Flux is energy/unit area

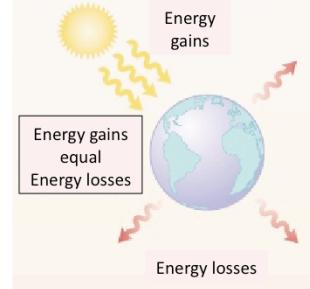
Not a coincidence that sun emits in our visible spectrum.

A consequence of evolution that our vision has evolved to optimize use of available spectrum



## Earth's Average Temperature

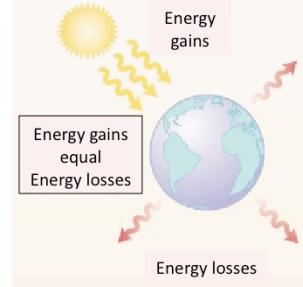
**Question:** What is Earth's average global annual temperature  $T$ ?  
What do you think?





## Earth's Average Temperature

**Question:** What is Earth's average global annual temperature  $T$ ?  
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**Answer:**  $15^{\circ}\text{C}$  ( $59^{\circ}\text{F}$ )

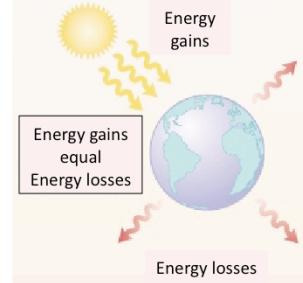




## Earth's Average Temperature

**Question:** What is Earth's average global annual temperature  $T_e$ ?  
What do you think?  
**Answer:**  $15^{\circ}\text{C}$  ( $59^{\circ}\text{F}$ )

**Question:** Why? What controls it?

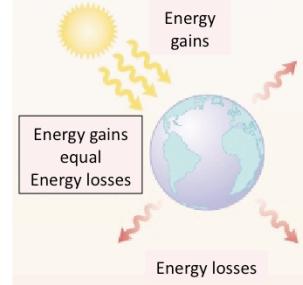




## Earth's Average Temperature

**Question:** What is Earth's average global annual temperature  $T_e$ ?  
What do you think?  
**Answer:** 15°C (59°F)

**Question:** Why? What controls it?  
**Answer:** Lots of things!





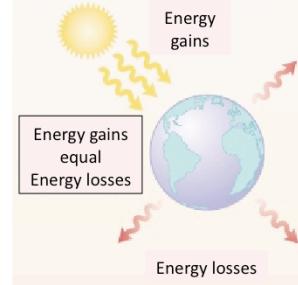
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**Let's use math to test this hypothesis:** Earth's temperature is a consequence of

1. Blackbody radiation
2. Distance from Sun
3. Size
4. Albedo (reflectivity)

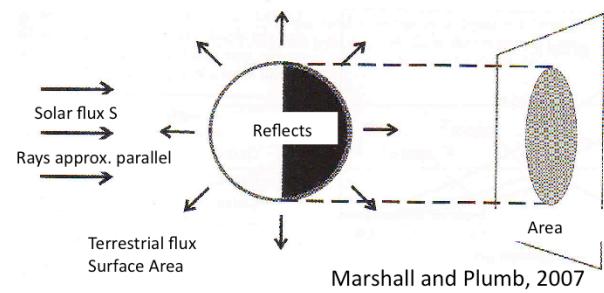
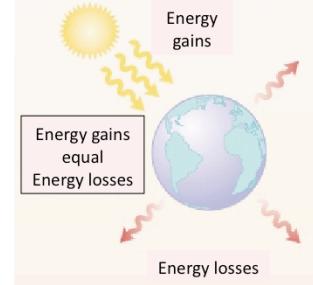
So we pretend the atmosphere is transparent to radiation: "no atmosphere"

Nice example of using mathematical models for hypothesis testing  
ATMOSPHERE TRANSPARENT to radiation means no greenhouse gases



## Energy Balance, No Atmosphere

- Hypothesis:** Earth's temperature is a consequence of
1. Blackbody radiation
  2. Distance from Sun
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Let's relate diagram carefully to each hypothesis

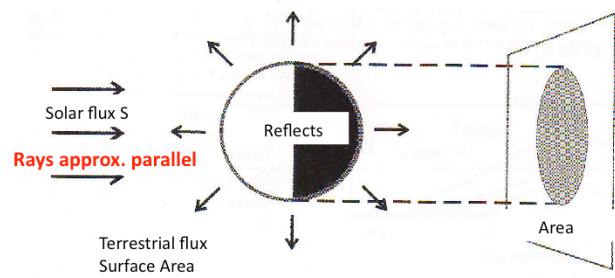
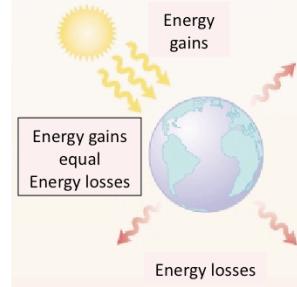


## Energy Balance, No Atmosphere

**Hypothesis:** Earth's temperature is a consequence of

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**Question:** Why are the rays from the sun approx. parallel? Which hypothesis?





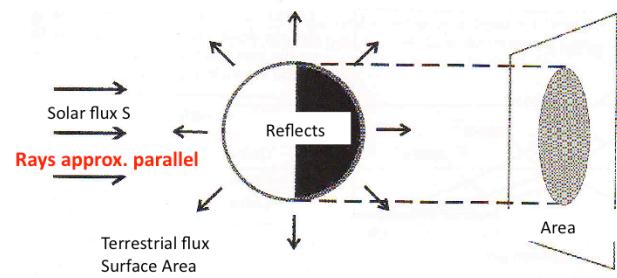
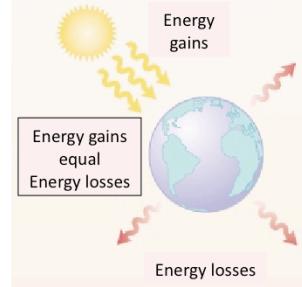
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**Question:** Why are the rays from the sun approx. parallel? Which hypothesis?

**Answer:** Sun is so far away (hypothesis 2)



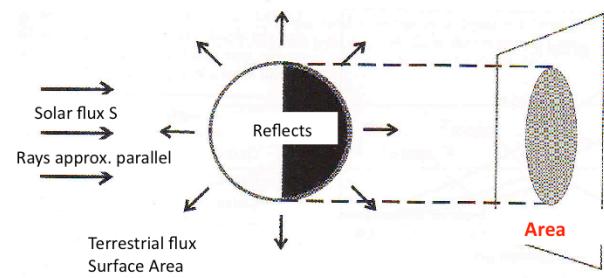
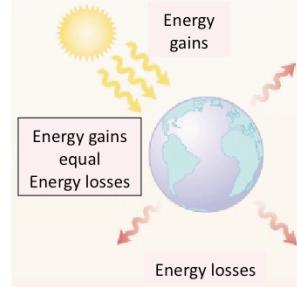


## Energy Balance, No Atmosphere

**Hypothesis:** Earth's temperature is a consequence of

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**Question:** What's the cross sectional area of the Earth (radius  $r$ )? Which hypothesis?

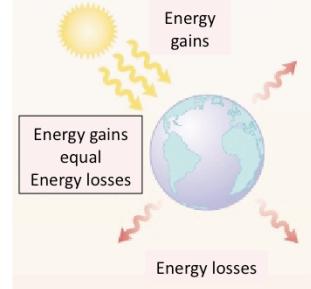




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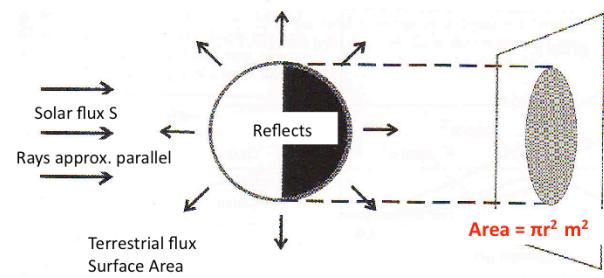
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**Question:** What's the cross sectional area of the Earth (radius  $r$ )? Which hypothesis?

**Answer:** Area =  $\pi r^2$  m<sup>2</sup> (hypothesis 3)



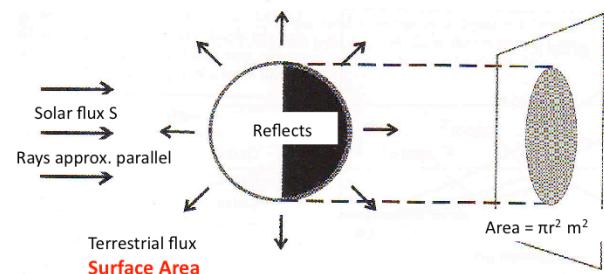
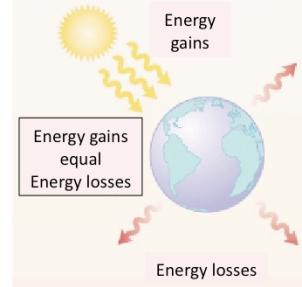


## Energy Balance, No Atmosphere

**Hypothesis:** Earth's temperature is a consequence of

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**Question:** What's the surface area of the Earth (radius  $r$ )? Which hypothesis?

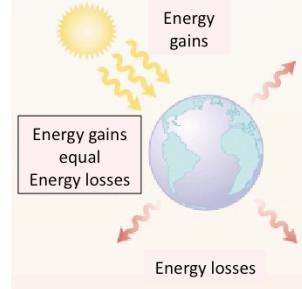




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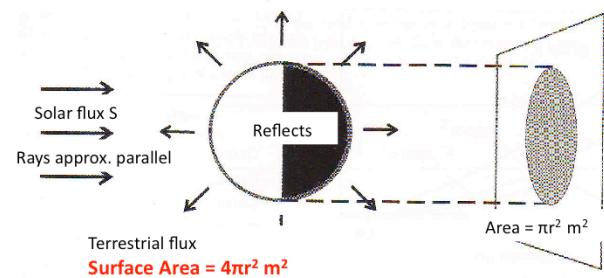
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**Question:** What's the surface area of the Earth (radius  $r$ )? Which hypothesis?

**Answer:** Surface Area =  $4\pi r^2$  (hypothesis 3)



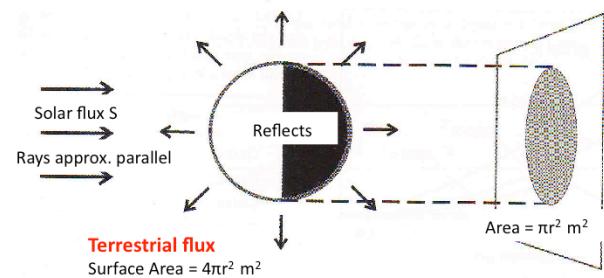
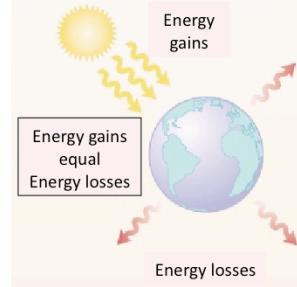


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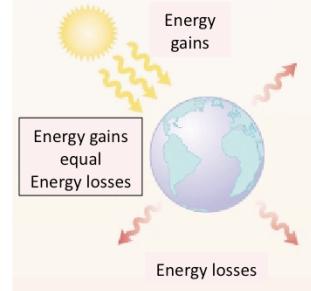
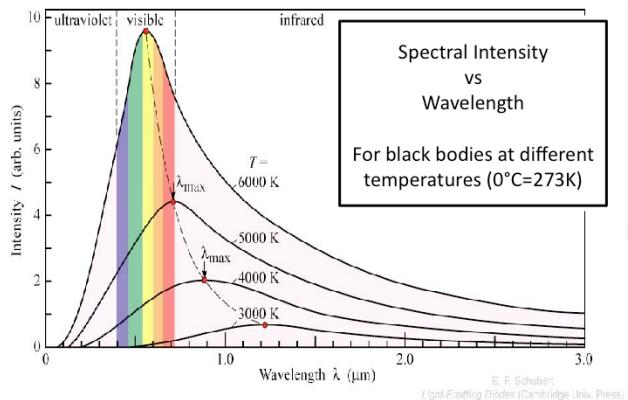
**Question:** What's the terrestrial flux? Which hypothesis?



This one is not so obvious – do you remember the formula?

# Figuring Out Terrestrial Flux

## Black Body Radiation - Planck



Sun: ~6000K, emits mainly in ultraviolet (UV) and visible spectrum (shortwave)  
 Earth: ~300K, emits mainly in infrared (IR) spectrum (longwave)

Total energy flux over all wavelengths for temp T is  $\sigma T^4 = (5.67 \times 10^{-8}) T^4$  -- Stefan-Boltzmann

We are going to use this to find an equation for T

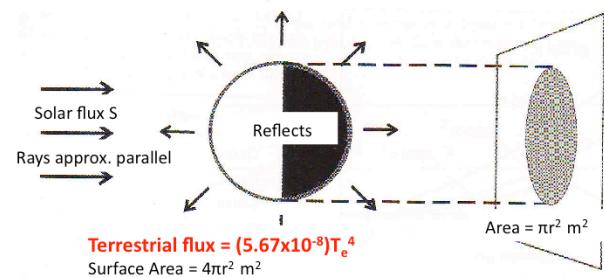
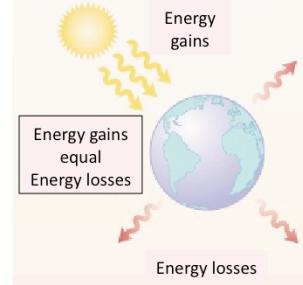


## Energy Balance, No Atmosphere

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**Question:** What's the terrestrial flux? Which hypothesis?

**Answer:** Terrestrial flux =  $(5.67 \times 10^{-8})T^4 \text{ W m}^{-2}$  (hyp. 1)



T is global annual average temperature of the earth

We already know it's 15 degrees because we've measured it, but we're going to build model equations to see what temp. these hypotheses predict.



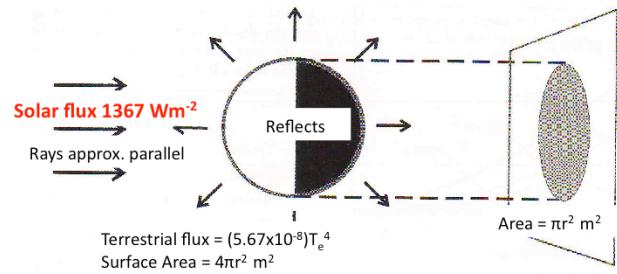
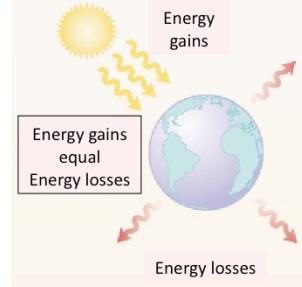
## Energy Balance, No Atmosphere

**Hypothesis:** Earth's temperature is a consequence of

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2. Distance from Sun
3. Size
4. Albedo (reflectivity)

**Question:** What's the Solar flux? Which hypothesis?

**Answer:** Solar flux  $S = 1367 \text{ Wm}^{-2}$  (hypothesis 1,2)



This is figured out using  $\sigma T^4$  for the sun, and the distance from the sun. Intensity drops like inverse square of distance.

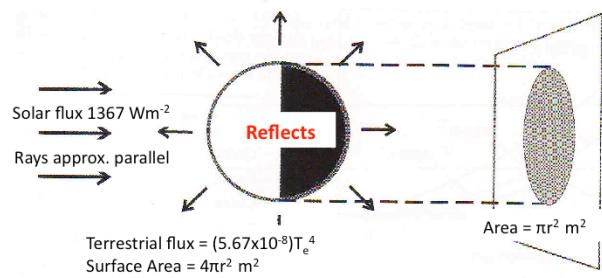
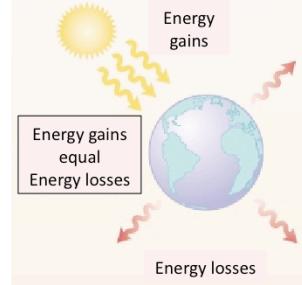


## Energy Balance, No Atmosphere

**Hypothesis:** Earth's temperature is a consequence of

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**Question:** What fraction of solar radiation does the earth reflect? What hypothesis?



What do you guess? On average over the whole planet?

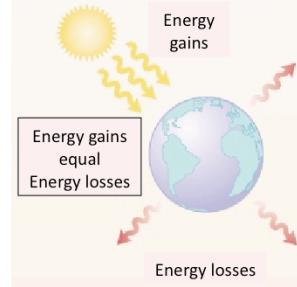


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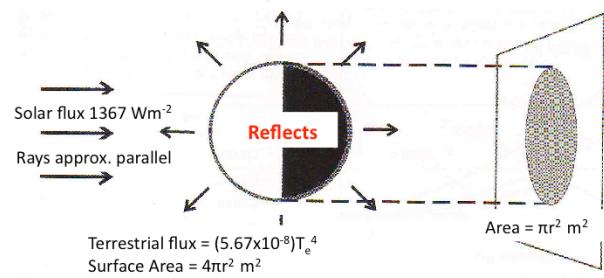
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**Question:** What fraction of solar radiation does the earth reflect? What hypothesis?



**Examples:**

- Fresh snow reflects 85%
- Sandy desert reflects 40%
- Grasslands reflect 18%
- Cities reflect 16%
- Ocean reflects 6%



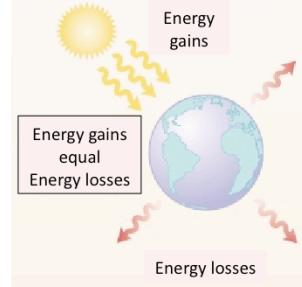
Albedo means the fraction that gets reflected



## Energy Balance, No Atmosphere

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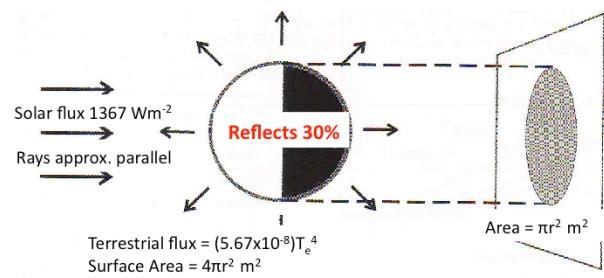


**Question:** What fraction of solar radiation does the earth reflect? What hypothesis?

**Answer:** Planetary average albedo 0.3 (hyp. 4)

**Examples:**

- Fresh snow reflects 85%
- Sandy desert reflects 40%
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- Ocean reflects 6%



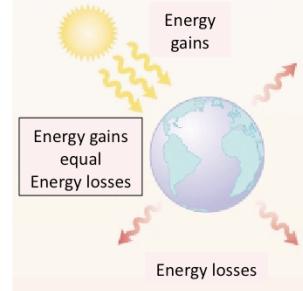
$$\begin{aligned} \text{Terrestrial flux} &= (5.67 \times 10^{-8}) T_e^4 \\ \text{Surface Area} &= 4\pi r^2 \text{ m}^2 \end{aligned}$$



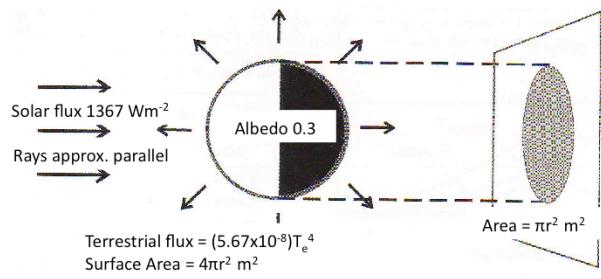
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OK! Now we understand the picture,  
Let's do energy balance...





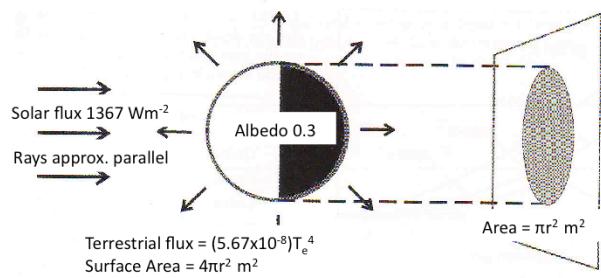
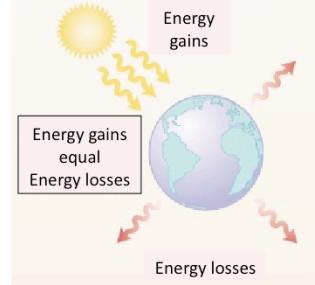
## Energy Balance, No Atmosphere

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### Energy Gain

Solar flux \* cross-sect area = ???



This is how much solar flux the earth intercepts, in total.

Flux is energy/unit area

So energy = flux \* area



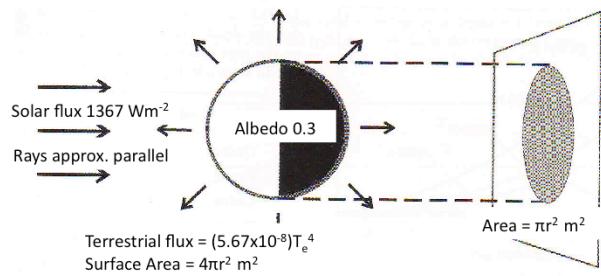
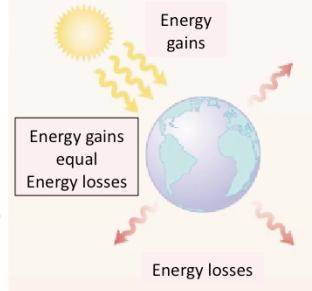
## Energy Balance, No Atmosphere

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### Energy Gain

$$\text{Solar flux} * \text{cross-sect area} = 1367 \text{ Wm}^{-2} * \pi r^2 \text{ m}^2 = 1367\pi r^2 \text{ W}$$





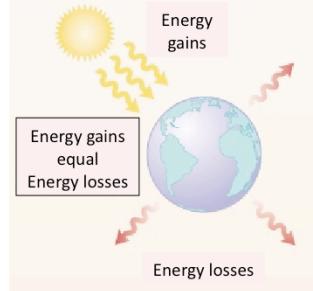
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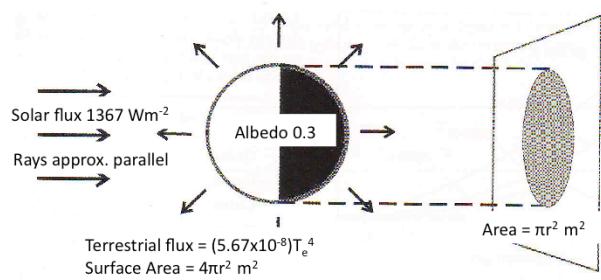
### Energy Gain

$$\text{Solar flux} * \text{cross-sect area} = 1367\pi r^2 \text{ W}$$



### Energy loss by Terrestrial Radiation

$$\text{Terrestrial flux} * \text{surface area} = ???$$



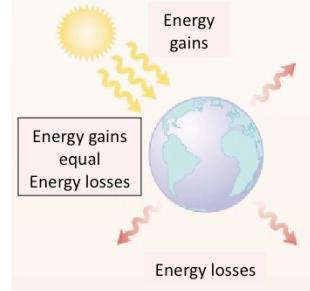


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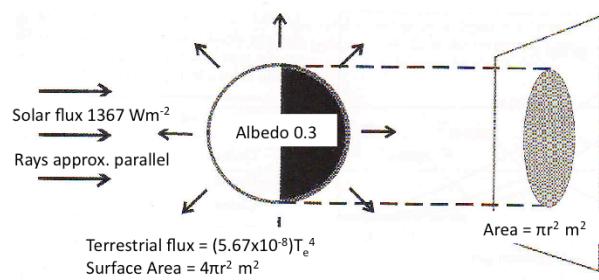
### Energy Gain

$$\text{Solar flux} * \text{cross-sect area} = 1367\pi r^2 \text{ W}$$



### Energy loss by Terrestrial Radiation

$$\text{Terrestrial flux} * \text{surface area} = (5.67 \times 10^{-8}) T^4 4\pi r^2 \text{ W}$$



Terrestrial flux given by Stefan-Bolzmann law of radiation from a black body at temperature T

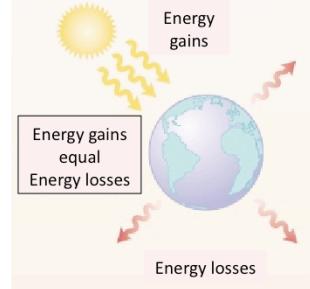


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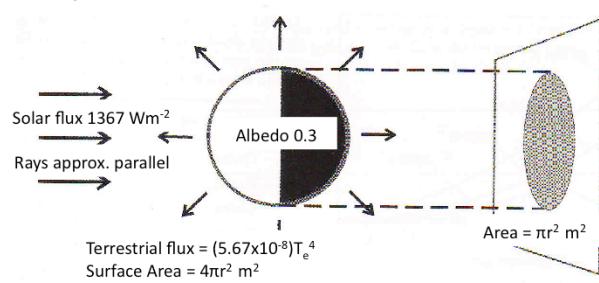


### Energy loss by Terrestrial Radiation

$$\text{Terrestrial flux} * \text{surface area} = (5.67 \times 10^{-8}) T^4 4\pi r^2 \text{ W}$$

### Energy loss by Albedo

Reflected solar radiation = ???



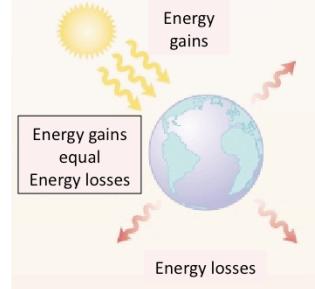
Albedo (reflectivity) \* solar radiation arriving



## Energy Balance, No Atmosphere

**Hypothesis:** Earth's temperature is a consequence of

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### Energy Gain

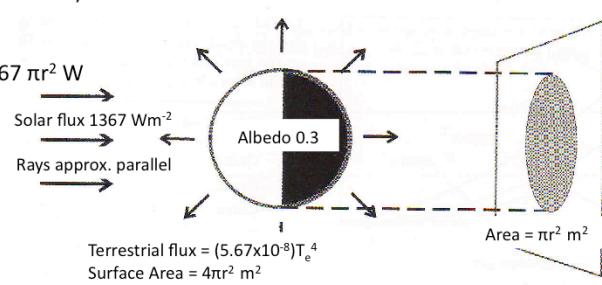
$$\text{Solar flux} * \text{cross-sect area} = 1367\pi r^2 \text{ W}$$

### Energy loss by Terrestrial Radiation

$$\text{Terrestrial flux} * \text{surface area} = (5.67 \times 10^{-8}) T^4 4\pi r^2 \text{ W}$$

### Energy loss by Albedo

$$\text{Reflected solar radiation} = 0.3 \times 1367 \pi r^2 \text{ W}$$

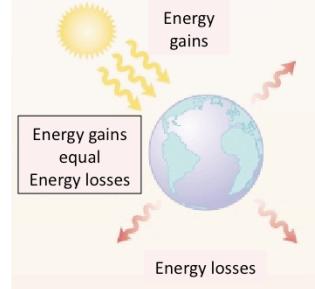




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### Energy Gain

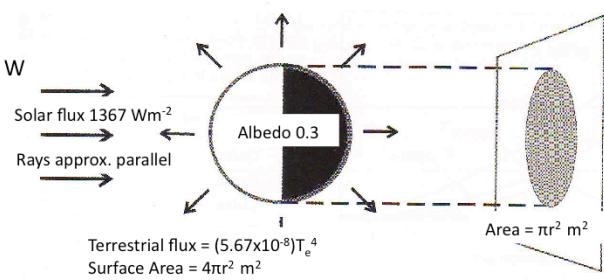
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### Energy loss by Terrestrial Radiation

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## Energy Balance, No Atmosphere

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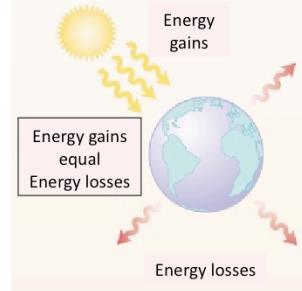
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### Balance

$$1367\pi r^2 = 410\pi r^2 + (5.67 \times 10^{-8})T^4 4\pi r^2$$



Cancel  $\pi r^2$  and solve for  $T$



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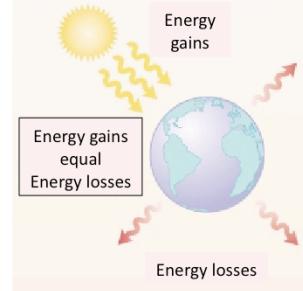
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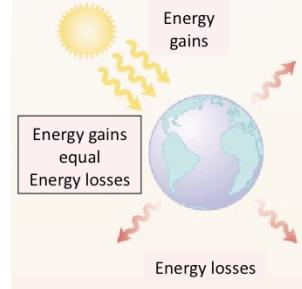
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$$(5.67 \times 10^{-8})T^4 = 239$$





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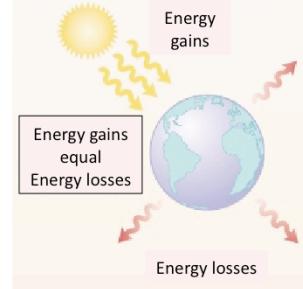
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$$T = 239^{(1/4)}(5.67 \times 10^{-8})^{(-1/4)}$$





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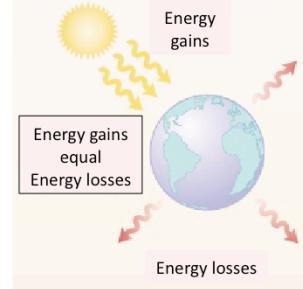
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$$T = 3.93 \times 64.8$$





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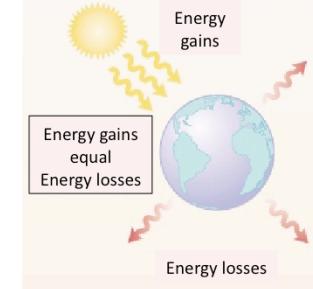
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So.....

$$T = 3.93 \times 64.8 = 255K$$

### Balance

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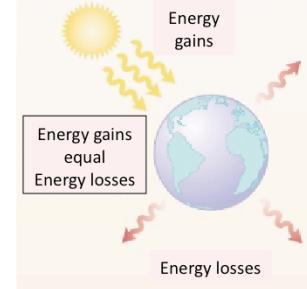
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So.....

$$T = 3.93 \times 64.8 = 255 \text{ K} = ???^\circ \text{C}$$

(Remember 273K = 0°C)



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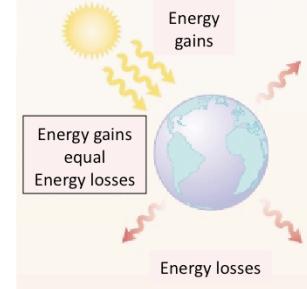
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So.....

$$T = 3.93 \times 64.8 = 255K = -18^{\circ}C$$

Does that sound about right?



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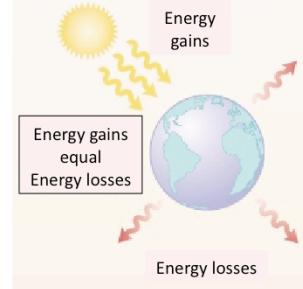
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Too Cold – Snowball Earth!



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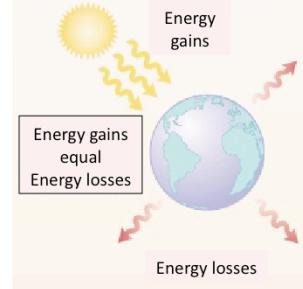
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Recall: our current global average surface temp is 15°C



## Energy Balance, No Atmosphere

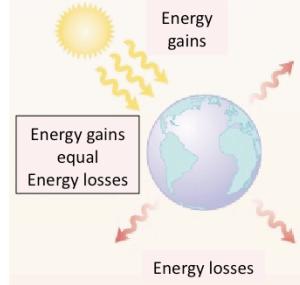
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**Conclusions:**

With these hypotheses, surface temp is 33°C too low.  
We need to include atmosphere to correct this.

255K is a good estimate for the top of the atmosphere,  
Which does satisfy the hypotheses more closely.



So.....

$$T = 3.93 \times 64.8 = 255K = -18^{\circ}C$$

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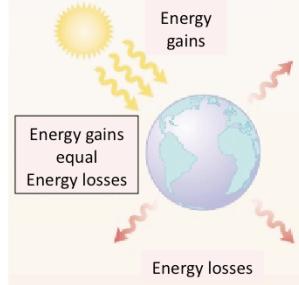
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Need new hypotheses about absorption of solar and  
terrestrial radiation by the atmosphere.



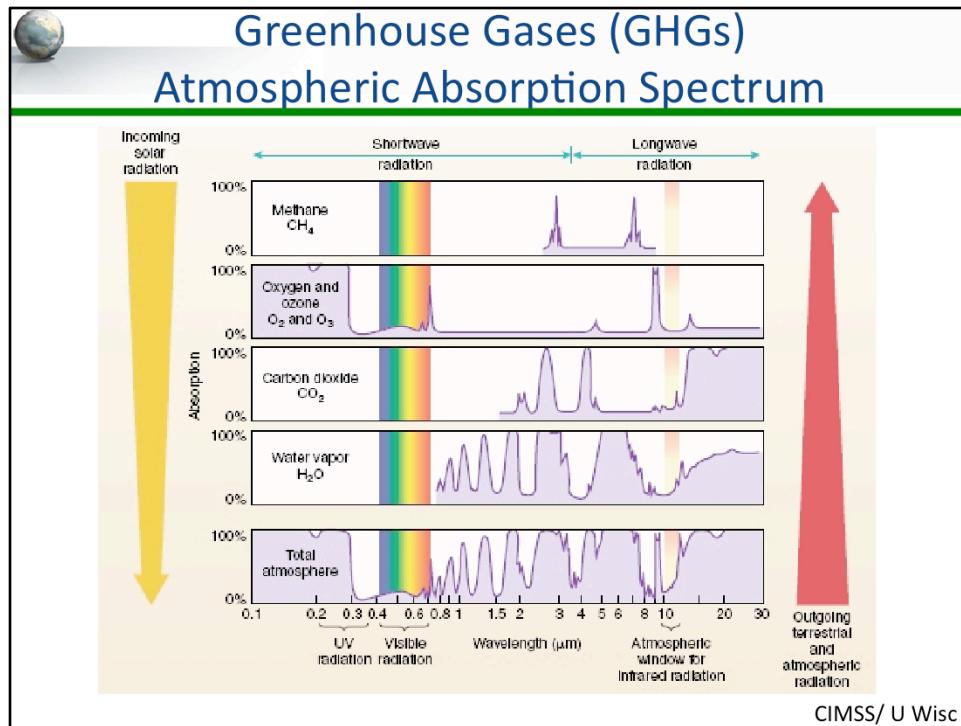
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CIMSS/ U Wisc

Atmosphere is not transparent.

Figure shows absorption spectrum due to major greenhouse gas constituents of the atmosphere

Shortwave range: solar radiation coming in – atmosphere is mainly transparent to that

Longwave (IR) range: terrestrial radiation leaving – atmosphere is partially opaque to IR

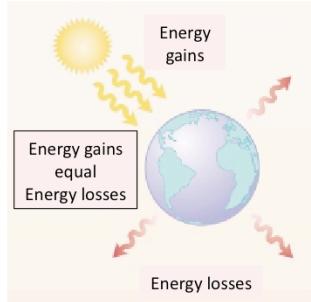


## Energy Balance, GHGs in Atmosphere

**Hypotheses:** Earth's temperature is a consequence of

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**New Balance**

$$1367\pi r^2 = 410\pi r^2 + \alpha(5.67 \times 10^{-8} T^4 4\pi r^2)$$

$$T = \frac{255}{\alpha^{1/4}}$$



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**Case  $\alpha=1$ :** Corresponds to no greenhouse gases

Atmosphere is transparent to terr. radiation, &

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Snowball Earth! We need *some* GHGs to avoid this.

As greenhouse gases increase:  $\alpha$  does what??? Increase or decrease?



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**So predicted temperature T does what ??? Increase or decrease?**



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Want to learn  
more???

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## To Learn More...



[www.mathclimate.org](http://www.mathclimate.org)

Mathematics and Climate Research Network – Education pages

[www.mpe2013.org](http://www.mpe2013.org)

Mathematics of Planet Earth 2013 – Education pages



[www.mathaware.org](http://www.mathaware.org)

Math Awareness Month 2013 – Mathematics and Sustainability

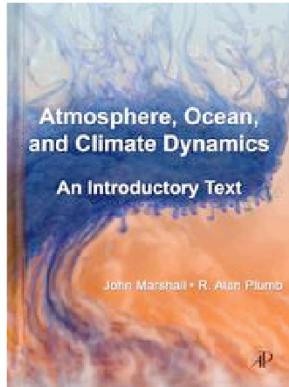
[www.dimacs.rutgers.edu/MPE](http://www.dimacs.rutgers.edu/MPE)

DIMACS Mathematics and Sustainability Curriculum Modules

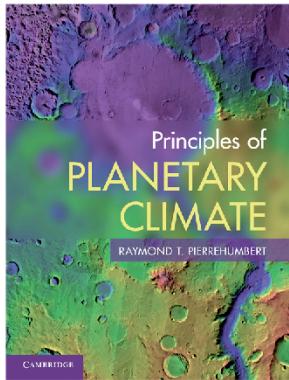




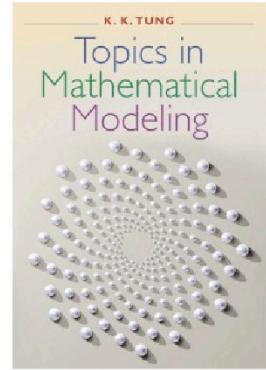
## And More...



John Marshall and  
Alan Plumb (MIT)



Ray Pierrehumbert



Ka-Kit Tung

Forthcoming text on Mathematics and Climate, Hans Kaper and Hans Engler. SIAM



## Math and Global Temperature

THANK  
YOU!



Thanks to:  
many friends  
& colleagues



& Bowdoin